ILLUMINATION GLOBALE

THEORIE ET PRATIQUE

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LIGHT

- Light = mixture of electromagnetic waves
- Each wave: frequency, period, wavelength, energy
- Wave called *spectral component*
- Radiation: emission or transport of light energy through a medium
- Each spectral component has a color appearance
- Light described by the amount of power in each of its spectral wavelength components
- Description : spectral power distribution $\Phi(\lambda)$ (SPD)
- Sampling $\Phi(\lambda)$
- Visible spectrum : [380nm, 780nm]
- Number of samples : 31 samples if wavelength spacing is 10nm
- We will see : 4 or 10 not equally spaced samples are sufficient
- SPD of the mixture of two lights = sum of the SPD's of the individual lights

• Light power Φ and its SPD $\Phi(\lambda)$ (Radiometry):

$$\Phi = \int_{380nm}^{780nm} \Phi(\lambda) d\lambda.$$

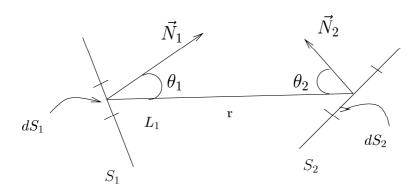
• Energetic and Luminous light powers (Phorometry):

$$\Phi_e = \int_0^\infty \Phi_e(\lambda) d\lambda
\Phi_v(\lambda) = 680 V(\lambda) \Phi_e(\lambda)
\Phi_v = \int_0^\infty \Phi_v(\lambda) d\lambda,$$

• $V(\lambda)$ is the sensitivity function : null out of [380nm, 780nm]

RADIOMETRY and PHOTOMETRY

- Radiometry: Measurement of quantities referring to radiation.
- **Photometry:** Measurement of quantities referring to radiation as evaluated according to a given luminous efficiency function, e.g. $V(\lambda)$.
- In the following, the expressions of all radiometric quantities are valid for light powers (Φ) as well as for each spectral component $(\Phi(\lambda))$.



- **light power or flux**: is the energy leaving a surface or impinging onto a surface per unit time.
- radiant intensity: is the flux leaving a surface per unit solid angle:

$$I = \frac{d\Phi}{d\Omega_1}$$

• radiance: is the flux leaving a surface per unit projected surface and per unit solid angle.

$$L_1 = \frac{d^2\Phi}{\cos\theta_1 dS_1 d\Omega_1}$$

where

$$d\Omega_1 = \frac{\cos \theta_2 dS_2}{r^2}$$

• radiant exitance: is also called radiant emittance or radiosity. It represents the light power leaving a surface, per unit area and is given by

$$B = \frac{d^2\Phi}{dS_1} = L_1 \cos \theta_1 d\Omega_1$$

.

• **irradiance**: is the light power, per unit area, inpinging onto a surface. It is expressed as

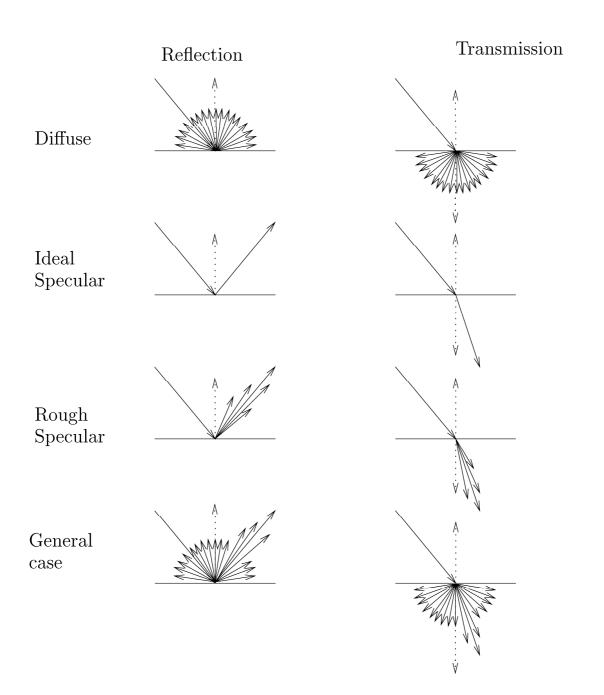
$$A = \frac{d^2\Phi}{dS_2} = L_1 \cos \theta_2 d\Omega_2$$

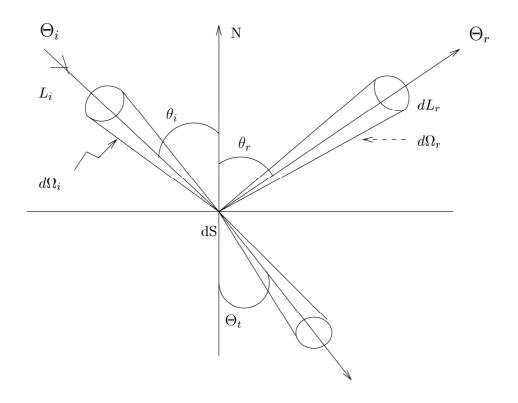
where

$$d\Omega_2 = \frac{dS_1 \cos \theta_1}{r^2}.$$

The following table recalls all these radiometric quantities.

	energetic quantity	luminous quantity
Flux	$\phi_e(\lambda) \; (Watt)$	$\phi_v(\lambda) \; (lumen)$
Radiant intensity	$I_e(\lambda) = \frac{d\phi_e(\lambda)}{d\Omega} \ (Watt.strd^{-1})$	$I_v \ (candela)$
Irradiance or Exitance	$E_e(\lambda) = \frac{d\phi_e(\lambda)}{dA} \ (Watt.m^{-2})$	$E_v \; (lux)$
Radiance	$L_e(\lambda) = \frac{d\phi_e(\lambda)}{d\omega dS \cos \alpha} (W.strd^{-1}.m^{-2})$	$L_v \; (cand.m^{-2})$





• Reflectivity: reflected power / incident power

$$\rho = \frac{d^2\Phi_r}{d^2\Phi_i} = \frac{B_r}{A_i} = \frac{radiosity}{irradiance}$$

• Bidirectional reflectance :

$$f_r(\Theta_i, \Theta_r) = \frac{dL_r}{dA} = \frac{radiance}{irradiance}.$$

• Or

$$f_r(\Theta_i, \Theta_r) = \frac{dL_r(\Theta_r)dS}{d^2\Phi_i(\Theta_i)}.$$

• If $d^2\Phi_i(\Theta_i)$ comes from a small light emitting surface of radiance L_i :

$$f_r(\Theta_i, \Theta_r) = \frac{dL_r}{L_i \cos \theta_i d\Omega_i}.$$

• If $d^3\Phi_r$ is the reflected power in direction Θ_r :

$$\frac{d^3\Phi_r}{d^2\Phi_i} = f_r(\Theta_i, \Theta_r) \cos \theta_r d\Omega_r.$$

• Thus:

$$\rho(\Theta_i) = \frac{d^2 \Phi_r}{d^2 \Phi_i} = \int_{2\pi} f_r(\Theta_i, \Theta_r) \cos \theta_r d\Omega_r.$$

• Bidirectional transmission : radiance dL_t in a direction of refraction Φ_t / irradiance for a direction of incidence Φ_i :

$$f_t(\Theta_i, \Theta_t) = \frac{dL_t}{dA} = \frac{radiance}{irradiance}.$$

Reflection Models

- Several physics based reflection models (COOK82,HE91,WARD92)
- Focus on Cook and Torrance model
- Reflected light depends on : wavelength, incidence angle, roughness, refractive index
- Polarization of light, masking shadowing of materials
- Surface approximated by small microfacets which are assumed to be perfectly specular
- The bidirectional reflectance:

$$f_r = sR_s + dR_d$$
 with $s + d = 1$

• R_d and R_s are respectively the diffuse and specular components, d and s are the proportions of the incident light which give rise to the diffuse and specular components respectively.

- $R_d = \frac{F(\lambda,0)}{\pi}$
- R_s accounts for roughness and masking/shadowing effects:

$$R_s = \frac{1}{4\pi} \frac{F(\lambda, \theta).D.G}{\cos \theta_i \cos \theta_r},$$

- Where:
 - $-F(\lambda,\theta)$ is the Fresnel factor
 - $-\theta_i$ is the incidence angle (direction D_i)
 - $-\theta_r$ the reflection angle (direction D_r)
 - $-\theta$ equals half of the angle (D_i, D_r)
 - -G accounts for the masking/shadowing effects between microfacets
 - -D characterizes the roughness of a surface (Beckman function)

Roughness

• D: microfacets distribution

$$D = \frac{1}{m^2 \cos^4 N \bullet H} e^{-[(\tan N \bullet H)/m]^2}$$

• G: masking and shadowing

$$G = \min\left(1, \frac{2(N \bullet H)(N \bullet V)}{(V \bullet H)}, \frac{2(N \bullet H)(N \bullet L)}{(V \bullet H)}\right)$$

Fresnel factor calculation

- We can find in books, for several materials, Fresnel factor curves $F(\lambda, 0)$ for normal incidence, as well as the refraction index \hat{n} for the wavelength $\tilde{\lambda} = 589$ (Sodium D lines) which corresponds to the center of the visible spectrum.
- Given these data, $F(\lambda, \theta)$ can be approximated, for each wavelength, by:

$$F(\lambda_i, \theta) = F(\lambda_i, 0)$$

$$+ \left(F(\lambda_i, \frac{\pi}{2}) - F(\lambda_i, 0) \right) \frac{F(\tilde{\lambda}, \theta) - F(\tilde{\lambda}, 0)}{F(\tilde{\lambda}, \frac{\pi}{2}) - F(\tilde{\lambda}, 0)},$$

where $F(\tilde{\lambda}, \theta)$ is given by the Fresnel formula for \hat{n} .

- If the values of the refraction index are given for a certain number of wavelengths, then compute exactly $F(\lambda_i, \theta)$ with the help of Fresnel formula.
- Knowing the expression of $F(\lambda, \theta)$, we can precompute it for each sample wavelength and for different values of θ (20 seem enough). These values allow to create a look-up table, from which any $F(\lambda, \theta)$ can be computed by a simple linear interpolation.

Transmission model

- So far, no physics-based transmission models have been proposed in the literature, but only an empirical one (HALL83).
- Rather than using an empirical transmission model, it is more realistic, for each material, to use transmittance values experimentally obtained with the help of a spectrophotometer.
- In case of ideal specular refraction, R_s is no more than $1 F(\lambda, \theta)$, and s = 1.

Fresnel Formula

 $F(\lambda, \theta_i)$: Fresnel Factor magnitude of reflected wave (Maxwell)

$$\bar{\mathcal{R}} = \frac{1}{2}(\mathcal{R}_{\parallel} + \mathcal{R}_{\perp})$$

• non metallic materials

$$\mathcal{R}_{\parallel} = \left(\frac{n\cos\theta_i - \cos\theta_t}{n\cos\theta_i + \cos\theta_t}\right)^2$$

$$\mathcal{R}_{\perp} = \left(\frac{\cos\theta_i - n\cos\theta_t}{n\cos\theta_i + n\cos\theta_t}\right)^2$$

• metallic materials

$$\mathcal{R}_{\parallel} = \frac{n^2(1+k^2)\cos^2\theta_i - 2n\rho\cos\theta_i(\cos\gamma + k\sin\gamma) + \rho^2}{n^2(1+k^2)\cos^2\theta_i + 2n\rho\cos\theta_i(\cos\gamma + k\sin\gamma) + \rho^2}$$

$$\mathcal{R}_{\perp} = \frac{\cos^2\theta_i + 2n\rho\cos\theta_i(k\sin\gamma - \cos\gamma) + n^2\rho^2(1+k^2)}{\cos^2\theta_i - 2n\rho\cos\theta_i(k\sin\gamma - \cos\gamma) + n^2\rho^2(1+k^2)}$$
with

$$\rho = \sqrt{A^2 + B^2}, \ \gamma = \frac{1}{2} \arctan \frac{B}{A},$$

$$A = 1 - \frac{(1-k^2)}{n^2(1+k^2)^2}\sin^2\theta_i, \ B = \frac{2k}{n^2(1+k^2)^2}\sin^2\theta_i$$

- Colorimetry: the science of measuring color based on the physical properties of light and the psychovisual properties of the human visual system.
- Maxwell tried to generate a large set of colors, by mixing three standard lights called color primaries.
- Result: most of the colors of the visible spectrum could be reproduced by combining only the three color primaries: Red, Green, Blue.
- These three color primaries must be linearly independent.
- They correspond to three different wavelengths.
- For a monitor, these three color primaries do not correspond to monochromatic lights.
- They act as a basis of a vector space called also *color* space.
- The coordinates of a color in this space are called trichromatric components or tristimulus values.

• The trichromatic components P_i of a light of spectral distribution $E(\lambda)$ is given by :

$$P_i = \int_{380nm}^{780nm} E(\lambda) \sigma_i(\lambda) d\lambda,$$

where the $\sigma_i(\lambda)$'s are called matching functions.

- RGB (CIE)
 - The Commission Internationale de l'Eclairage (CIE) proposed in 1930, three color primaries: Red, Green and Blue.
 - In case of the RGB color space of a monitor, the three associated matching functions are $\bar{r}, \bar{g}, \bar{b}$. They depend on the used monitor.

CIE XYZ color space

- The CIE has normalised a color space, in which the three color primaries X, Y, and Z are not physical colors.
- The advantage of this color space is that it is independent of the used display device.
- The particularity of this color space is that the Y component corresponds to the visual luminance of the spectrum and is obtained by taking into account the sensitivity of a reference observer $(\bar{y}(\lambda) = V(\lambda))$.

• Trichromatic components X,Y and Z:

$$X = K \int_{380}^{780} E(\lambda) \bar{x}(\lambda) d\lambda$$

$$Y = K \int_{380}^{780} E(\lambda) \bar{y}(\lambda) d\lambda$$

$$Z = K \int_{380}^{780} E(\lambda) \bar{z}(\lambda) d\lambda$$

- For an absolute SPD, $K = 680 \, lumen/watt$.
- For a relative SPD, K is selected such that bright white has a Y value of 100, then other Y values will be in the range of 0 to 100.
- Thus:

$$K = 100 / \int E_w(\lambda) \bar{y}(\lambda) d\lambda$$

where $E_w(\lambda)$ is the SPD for any standard white light source (D6500).

• The CIE standard chromaticity coordinates x, y, z are generated by projecting the tristimulus values on the X + Y + Z = 1 plane so that:

$$x = X / (X + Y + Z)$$

 $y = Y / (X + Y + Z)$
 $z = Z / (X + Y + Z)$
 $1 = x + y + z$

- A common specification for color is : Y, x, y, where Y describes the luminance of the color (response to brightness) and x, y defines a point on the chromaticity diagram.
- The chromaticity diagram gives an indication of the color independent of its brightness.
- The CIE chromaticity diagram is widely used in industry for describing colors.

Transformation from XYZ to RGB

• The transformation of a color from space RGB to space XYZ is expressed as:

$$\begin{pmatrix} X \\ Y \\ Z \end{pmatrix} = \begin{pmatrix} X_r & X_g & X_b \\ Y_r & Y_g & Y_b \\ Z_r & Z_g & Z_b \end{pmatrix} \begin{pmatrix} R \\ G \\ B \end{pmatrix}$$

- The coefficients of the transformation matrix depend on :
 - the chromaticity coordinates $x_r, y_r, x_g, y_g, x_b, y_b$ of the phosphors of the display device,
 - the chromaticity coordinates x_w et y_w of the white point of the display device,
 - the luminance Y_w of this white point.

Chromatic distance between colors

- CIELUV space

- This color space (also known as $L^*u^*v^*$), established in 1964 and adopted by the CIE in 1978.
- The three components in this space are expressed by:

$$L^* = 166(Y/Y_n)^{0.5} - 16, Y/Y_n > 0.01,$$

$$u^* = 13L^*(u' - u_n)$$

$$v^* = 13L^*(v' - v_n)$$

where

$$u' = 4X/(X + 15Y + 3Z)$$

$$v' = 9Y/(X + 15Y + 3Z)$$

$$u_n = 4X_n/(X_n + 15Y_n + 3Z_n)$$

$$v_n = 9Y_n/(X_n + 15Y_n + 3Z_n)$$

- X_n , Y_n and Y_n being the trichromatic components of the reference white (ex: D6500).
- In this space the difference between two colors is expressed as:

$$\Delta E = (\Delta L^{*2} + \Delta u^{*2} + \Delta v^{*2})^{0.5}$$

• Detection of small differences

Chromatic distance between colors

- CIELAB space

- Another system called CIELAB or $L^*a^*b^*$, more suitable for measuring important differences between colors.
- The difference is also expressed as:

$$\Delta E = (\Delta L^{*2} + \Delta a^{*2} + \Delta b^{*2})^{0.5}$$

• Where:

$$L^* = 166(Y/Y_n)^{0.33} - 16$$

$$a^* = 500[(X/X_n)^{0.33} - (Y/Y_n)^{0.33}]$$

$$b^* = 200[(Y/Y_n)^{0.33} - (Z/Z_n)^{0.33}]$$

if (X/X_n) , (Y/Y_n) and (Z/Z_n) are bigger than 0.01.

TRICHROMATIC APPROACH VERSUS SPECTRAL APPROACH

- Two approaches for computing a synthetic image: trichromatic and spectral.
- Spectral: considers spectra (spectral distribution of light, spectral reflectance, transmittance and absorption, refraction index depending on wavelength...) instead of trichromatic components.
 - If $E(\lambda)$ is the incoming light, and $f_r(\lambda)$ the bidirectional reflectance, then

$$S(\lambda) = f_r(\lambda) \times E(\lambda)$$

- The RGB components S_R , S_G and S_B of the reflected light are obtained by:

$$S_{R} = \int_{380}^{780} f_{r}(\lambda) \times E(\lambda) \, \bar{r}(\lambda) \, d\lambda$$

$$S_{V} = \int_{380}^{780} f_{r}(\lambda) \times E(\lambda) \, \bar{v}(\lambda) \, d\lambda$$

$$S_{B} = \int_{380}^{780} f_{r}(\lambda) \times E(\lambda) \, \bar{b}(\lambda) \, d\lambda \qquad (1)$$

- Trichromatic: the quantities must be described by their RGB components: E_R , E_G and E_B for the incident light, and f_r^R , f_r^G and f_r^B for the reflectance function.
 - The RGB components of the reflected light are:

$$S_R = f_r^R \times E_R$$

$$S_G = f_r^G \times E_G$$

$$S_B = f_r^B \times E_B$$

- These triplets are obtained by:

$$E_R = \int_{380}^{780} E(\lambda) \ \bar{r}(\lambda) \ d\lambda$$

$$E_G = \int_{380}^{780} E(\lambda) \ \bar{g}(\lambda) \ d\lambda$$

$$E_B = \int_{380}^{780} E(\lambda) \ \bar{b}(\lambda) \ d\lambda$$

$$f_r^R = \int_{380}^{780} f_r(\lambda) \ \bar{r}(\lambda) \ d\lambda$$

$$f_r^G = \int_{380}^{780} f_r(\lambda) \ \bar{g}(\lambda) \ d\lambda$$

$$f_r^B = \int_{380}^{780} f_r(\lambda) \ \bar{b}(\lambda) \ d\lambda.$$

Comparision of these two approaches

• The spectral approach leads to:

$$S_R = \int_{380}^{780} f_r(\lambda) E(\lambda) \ \bar{r}(\lambda) \ d\lambda,$$

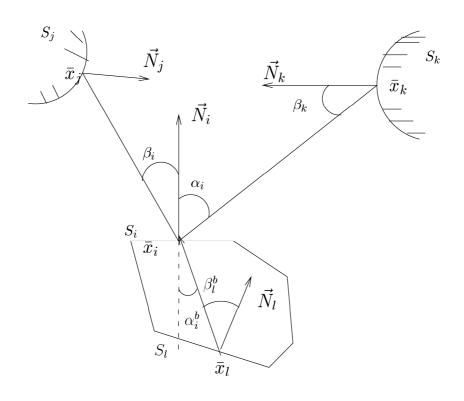
• The trichromatic approach gives:

$$S_R = \int_{380}^{780} f_r(\lambda) \ \bar{r}(\lambda) \ d\lambda \int_{380}^{780} E(\lambda) \ \bar{r}(\lambda) \ d\lambda$$

• These two last equations show that the trichromatic approach approximates an integral by the product of two integrals, which is not mathematically correct.

Spectral sampling

- To display a light on a display device, the three trichromatic components RGB of its spectral distribution have to be calculated.
- The accuracy of this calculation strongly depends on the way the visible spectrum is sampled. It depends on both the sample values and their number.
- Meyer's method: use The AC1C2 color space, its axes are oriented along the most dense color regions, each one having an importance that is proportional to the density of these regions.



- α_i and β_i refer, respectively, to angle of incidence and angle of reflection at point \bar{x}_i of a surface S_i .
- α_i^b and β_l^b refer, respectively, to angle of incidence on the back of surface S_i and angle of transmission at point \bar{x}_l .
- in all subscript or function argument notations, the order of the subscripts or the arguments follows the propagation of light with the source being the leftmost.

The global model

- $L(\bar{x}_i, \bar{x}_j)$: the radiance of surface S_i at point \bar{x}_i as seen from point \bar{x}_i at surface S_j .
- Summing the contributions of all surfaces S_k , we have:

$$L(\bar{x}_i, \bar{x}_j) = L^e(\bar{x}_i, \bar{x}_j) + \sum_k \int_{\Omega_{ik}} f_r(\bar{x}_k, \bar{x}_i, \bar{x}_j) L(\bar{x}_k, \bar{x}_i) \cos \alpha_i d\Omega_{ik}$$
$$+ \sum_l \int_{\Omega_{il}^b} e^{-\sigma \tau} f_t(\bar{x}_l, \bar{x}_i, \bar{x}_j) L(\bar{x}_l, \bar{x}_i) \cos \alpha_i^b d\Omega_{il}^b$$

- $f_r(\bar{x}_k, \bar{x}_i, \bar{x}_j)$ and $f_t(\bar{x}_l, \bar{x}_i, \bar{x}_j)$: bidirectional reflection and transmittance repectively.
- Ω_{ik} is the solid angle under which surface S_k is seen at point \bar{x}_i .
- Ω_{il}^b is the solid angle corresponding to the incident directions on the back of surface i, under which surface S_l is seen at point \bar{x}_i .
- $L^e(\bar{x}_i, \bar{x}_j)$ is the radiance due to self-emittance.
- τ the traveling length of the incident ray into the transparent object, say $\bar{x}_l\bar{x}_i$.

The global model

$$L(\bar{x}_{i}, \Theta_{out}) = L^{e}(\bar{x}_{i}, \Theta_{out}) + \int_{2\pi} f_{r}(\bar{x}_{i}, \Theta_{in}, \Theta_{out}) L(\bar{x}_{i}, \Theta_{in}) \cos \alpha_{i} d\Omega_{i}$$
$$+ \int_{\Omega_{i}^{b}} e^{-\sigma \tau} f_{r}(\bar{x}_{i}, \Theta_{in}^{b}, \Theta_{out}) L(\bar{x}_{i}, \Theta_{in}^{b}) \cos \alpha_{i}^{b} d\Omega_{i}^{b}$$

Reflexion only:

$$L(\bar{x}_i, \Theta_{out}) = L^e(\bar{x}_i, \Theta_{out}) + \int_{2\pi} f_r(\bar{x}_i, \Theta_{in}, \Theta_{out}) L(\bar{x}_i, \Theta_{in}) \cos \alpha_i d\Omega_i$$

The global model

• As

$$d\Omega_{ik} = \frac{dS_k \cos \beta_k}{||\bar{x}_k \bar{x}_i||^2}, \ d\Omega_{il}^b = \frac{dS_l \cos \beta_l^b}{||\bar{x}_l \bar{x}_i||^2}$$

• We have:

$$L(\bar{x}_{i}, \bar{x}_{j}) = L^{e}(\bar{x}_{i}, \bar{x}_{j}) + \sum_{k} \int_{S_{k}} f_{r}(\bar{x}_{k}, \bar{x}_{i}, \bar{x}_{j}) L(\bar{x}_{k}, \bar{x}_{i}) G(\bar{x}_{k}, \bar{x}_{i}) dS_{k}$$
$$+ \sum_{l} \int_{S_{l}} e^{-\sigma \tau} f_{t}(\bar{x}_{l}, \bar{x}_{i}, \bar{x}_{j}) L(\bar{x}_{l}, \bar{x}_{i}) G'(\bar{x}_{l}, \bar{x}_{i}) dS_{l}$$

• Or:

$$L(\bar{x}_{i}, \bar{x}_{j}) = L^{e}(\bar{x}_{i}, \bar{x}_{j}) + \sum_{k} \int_{\bar{x}_{k} \in S_{k}} f_{r}(\bar{x}_{k}, \bar{x}_{i}, \bar{x}_{j}) L(\bar{x}_{k}, \bar{x}_{i}) G(\bar{x}_{k}, \bar{x}_{i}) d\bar{x}_{k}$$
$$+ \sum_{l} \int_{\bar{x}_{l} \in S_{l}} e^{-\sigma\tau} f_{t}(\bar{x}_{l}, \bar{x}_{i}, \bar{x}_{j}) L(\bar{x}_{l}, \bar{x}_{i}) G'(\bar{x}_{l}, \bar{x}_{i}) d\bar{x}_{l}$$

• where $G(\bar{x_k}, \bar{x_i}), G'(\bar{x_l}, \bar{x_i})$ are purely geometric terms as

$$G(\bar{x_k}, \bar{x_i}) = \frac{\cos \alpha_i \cos \beta_k}{||\bar{x_k}\bar{x_i}||^2}, \ G'(\bar{x_l}, \bar{x_i}) = \frac{\cos \alpha_i^b \cos \beta_l^b}{||\bar{x_l}\bar{x_i}||^2}.$$

The global model

- The light occlusion effect can be accounted for by introducing a function $h(\bar{x}_i, \bar{x}_j)$ taking the value 1 if point \bar{x}_i is visible from point \bar{x}_j and 0 otherwise.
- Thus

$$L(\bar{x}_{i}, \bar{x}_{j}) = h(\bar{x}_{i}, \bar{x}_{j})[L^{e}(\bar{x}_{i}, \bar{x}_{j}) + \sum_{k} \int_{\bar{x}_{k} \in S_{k}} f_{r}(\bar{x}_{k}, \bar{x}_{i}, \bar{x}_{j}) L(\bar{x}_{k}, \bar{x}_{i}) G(\bar{x}_{k}, \bar{x}_{i}) d\bar{x}_{k} + \sum_{l} \int_{\bar{x}_{l} \in S_{l}} e^{-\sigma \tau} f_{t}(\bar{x}_{l}, \bar{x}_{i}, \bar{x}_{j}) L(\bar{x}_{l}, \bar{x}_{i}) G'(\bar{x}_{l}, \bar{x}_{i}) d\bar{x}_{l}]$$

- The above system of equations completely describe the light transport mechanisms between surfaces.
- The knowledge of $L(\bar{x}_i, \bar{x}_j)$ is sufficient to describe the spatial distribution of the light radiating from surface S_i .

RADIOSITY

- Surfaces perfectly diffuse.
- Thus

$$f_r(\bar{x}_k, \bar{x}_i, \bar{x}_j) = R^d(\bar{x}_i), d = 1$$
$$L(\bar{x}_i, \bar{x}_j) = L(\bar{x}_i)$$

• And

$$L(\bar{x}_{i}) = L^{e}(\bar{x}_{i}) + R^{d}(\bar{x}_{i}) \sum_{k} \int_{\bar{x}_{k} \in S_{k}} L(\bar{x}_{k}, \bar{x}_{i}) G(\bar{x}_{k}, \bar{x}_{i}) h(\bar{x}_{k}, \bar{x}_{i}) d\bar{x}_{k}$$

• Or:

$$L(\bar{x}) = L^{e}(\bar{x}) + R^{d}(\bar{x}) \int_{\bar{y} \in S} L(\bar{y}) G(\bar{y}, \bar{x}) h(\bar{y}, \bar{x}) d\bar{y}$$

• If we multiply by π :

$$B(\bar{x}_i) = E(\bar{x}_i)$$

$$+ R^d(\bar{x}_i) \sum_k \int_{\bar{x}_k \in S_k} B(\bar{x}_k) G(\bar{x}_k, \bar{x}_i) h(\bar{x}_k, \bar{x}_i) d\bar{x}_k$$

• Or:

$$\begin{array}{ll} B(\bar{x}) \ = \ E(\bar{x}) \\ + \ R^d(\bar{x}) \int_{\bar{y} \in S} B(\bar{y}) G(\bar{y}, \bar{x}) h(\bar{y}, \bar{x}) d\bar{y} \end{array}$$

• If all the surfaces are meshed so that $B(\bar{x}_i) = constant, \forall i$:

$$B_i = \frac{1}{S_i} \int_{\bar{x_i} \in S_i} B(\bar{x_i}) d\bar{x_i}$$

• Then we get:

$$B_{i} = E_{i} + \rho_{i} \sum_{k} F_{ik} B_{k}$$

$$A_{i}B_{i} = A_{i}E_{i} + \rho_{i} \sum_{k} F_{ik} A_{i} B_{k}$$

$$A_{i}B_{i} = A_{i}E_{i} + \rho_{i} \sum_{k} F_{ki} A_{k} B_{k}$$

$$\Phi_{i} = \Phi_{i}^{E} + \rho_{i} \sum_{k} F_{ki} \Phi_{k}$$

$$A_{i}F_{ik} = A_{k}F_{ki}, reciprocite$$

• Where

$$-R^{d}(\bar{x}) = \rho_i/\pi$$

$$F_{ik} = \frac{1}{\pi S_i} \int_{\bar{x_i} \in S_i} \int_{\bar{x_k} \in S_k} G(\bar{x_k}, \bar{x_i}) h(\bar{x_k}, \bar{x_i}) d\bar{x_k} d\bar{x_i}$$

 $-F_{ik}$ is called form factor.

RADIOSITY

The system of equations

$$B_i = E_i + \rho_i \sum_k F_{ik} B_k$$

- B_i : Exitance of patch i (Radiosity);
- E_i : self-emitted radiosity of patch i;
- ρ_i : reflectivity of patch i;
- F_{ik} : form-factor giving the fraction of the energy leaving patch i that arrives at patch k;
- \bullet N: number of patches.

CONSTANT RADIOSITY

The different steps

- Discretize the objects' scene into small patches.
- Calculate the form-factors, then the system matrix.
- Solve this system.
- Calculate the radiosity of each vertex of each patch by averaging the radiosities (either $B_i(\lambda)$ or B_R , B_G , B_G) of the patches sharing it. Divide them by π to convert them to radiances (if spectral approach,

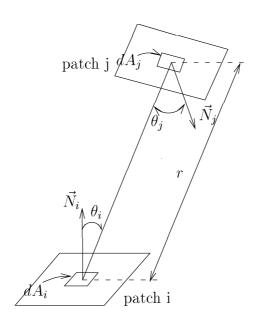
$$B_R = \int_{380}^{780} L_i(\lambda) \bar{r}(\lambda) d\lambda.$$

• Render the image by Z-buffering with Gouraud shading, or by ray tracing.

Remarks

- Solution independent of the viewpoint
- Thus, when moving the viewpoint, only the rendering step has to be run.
- Which can be handled in one second on specific graphics station. Interactivity

Form-factor calculation



• Expression :

$$F_{ij} = \frac{1}{\pi A_i} \int_{A_i} \int_{A_j} \frac{h(\bar{x}_i, \bar{x}_j) \cos \theta_i \cos \theta_j}{r^2} dA_j dA_i,$$

• The form factor between a differential element of patch i (around a point P_i) and patch j is:

$$F_{dA_iA_j} = \frac{1}{\pi} \int_{A_j} \frac{h(\bar{x}_i, \bar{x}_j) \cos \theta_i \cos \theta_j}{r^2} dA_j,$$

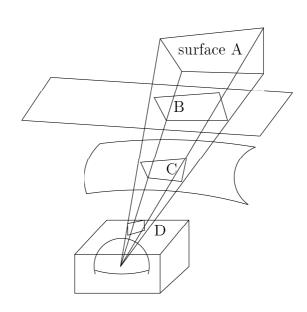
• If the two patches are far enough, this form factor is a good guess for F_{ij} .

- To compute F_{ij} , patch i is subdivided into R small elements dA_i^q and all $F_{dA_i^qA_j}$'s are evaluated.
- F_{ij} is then equal to:

$$F_{ij} = \frac{1}{A_i} \sum_{q=1}^R F_{dA_i^q A_j} dA_i^q$$

Projection methods

- If two patches similarly project on a given projection surface, then their form-factor (with a differential element of another patch) is thus similar.
- Find a suitable projection surface (Hemi-cube, Hemi-sphere, Plane) to simpify the form-factor calculation.

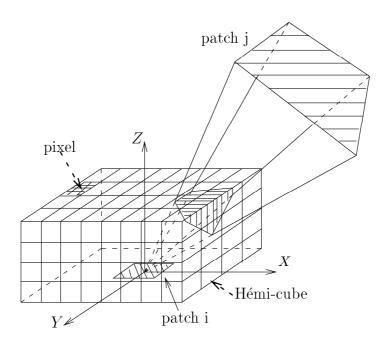


Surface Projection: HEMICUBE

Objective: compute the form factors between a patch i and all the other patches with the help of a projection surface.

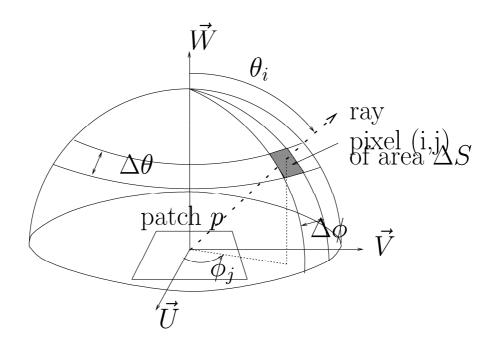
- Hemi-cube: imaginary half-cube placed at the center of the receiving patch element.
- A coordinate system is associated with this Hemicube, whose positive Z axis coincides with the normal to the Patch.
- Projection of the other patches onto the five faces of the Hemi-cube.
- Faces discretized into pixels

Surface Projection: HEMICUBE



- Transformation of the environment into this coordinate system.
- Clipping of the environment, Projection onto the five faces, scan-conversion.
- If two patches project on the same pixel of the hemicube, then Z-buffering.
- An item buffer is maintained, giving for each pixel the patch seen from the origin of the coordinate system.
- A delta form-factor is found for a differential element dA_i to a pixel and stored in a look-up table.
- After determining which patch A_j is visible at each pixel on the hemi-cube, a summation of the delta form-factors for each pixel occupied by patch A_j (item buffer) determines the form-factor from the patch element dA_i to patch A_j .
- Then the hemi-cube is placed around another differential element of patch A_i .
- Once all the differential elements of patch A_i have been considered, the form-factors F_{ij} are evaluated, and the hemi-cube is positioned at the center of a differential element of another patch.

Surface Projection: HEMISPHERE



- Use a hemisphere as a projection surface and ray tracing.
- A hemisphere is placed at the center of a patch p and is discretized by sampling the two polar angles θ and ϕ .
- The hemisphere is then discretized into surface elements ΔS , each one corresponds to a small form-factor called delta form-factor.
- Delta form-factor associated with a pixel:

$$\Delta FF = \frac{1}{\pi}\cos\theta\sin\theta\Delta\theta\Delta\phi$$

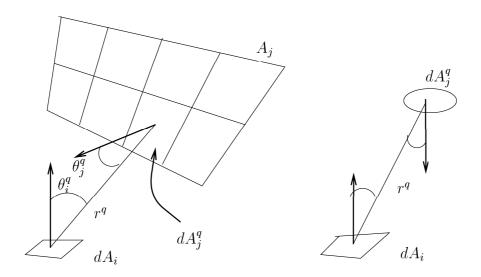
- Visibilty: a ray is cast from the hemisphere center through each pixel \Rightarrow intersection.
- The identifier of the patch containing the closest intersection point is stored in an item buffer.

- Once all the rays have been cast toward all directions (θ_i, ϕ_j) , the form-factors are calculated by scanning the item buffer and summing the delta form-factors associated with the rays along which a particular patch is visible.
- This form-factor calculation technique is simpler and faster than the hemicube approach, since it avoids several processings such as polygon clipping, polygon filling and geometric transformations.

Non projection techniques

ullet Goal: compute the form factors for each pair (i,j) of patches.

• Ray Tracing



• This modified point-to-disk formula is given by:

$$F_{dA_i dA_j^q} = \frac{dA_j^q \cos \theta_i^q \cos \theta_j^q}{\pi (r^q)^2 + dA_j^q}$$

• Then

$$F_{dA_{i}A_{j}} = \sum_{q=1}^{R} \frac{dA_{j}^{q} \cos \theta_{i}^{q} \cos \theta_{j}^{q}}{\pi (r^{q})^{2} + dA_{j}^{q}} h(dA_{i}, dA_{j}^{q}),$$

ullet The point-to-disk formula breaks down if the distance r is small relative to the differential area.

Monte Carlo method

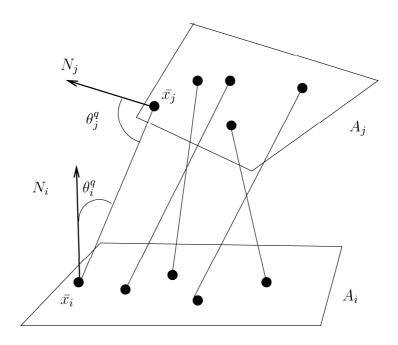


Figure 1: area-area form factor calculated by Monte Carlo

Pseudo code for Monte Carlo area-to-area form factor computation

```
Fij = 0

for k = 1 to n do

randomly select point \bar{x}_i on the element i

randomly select point \bar{x}_j on the element j

determine visibility between \bar{x}_i and \bar{x}_j

if visible

compute r^2 = (\bar{x}_i - \bar{x}_j)^2

compute \cos \theta_i = \vec{r}_{ij} \bullet \vec{N}_i

compute \cos \theta_j = \vec{r}_{ji} \bullet \vec{N}_j

compute \Delta F = \frac{\cos \theta_i \cos \theta_j}{\pi r^2 + \frac{A_j}{n}}

if (\Delta F > 0) F_{ij} = F_{ij} + \Delta F

F_{ij} = F_{ij} * A_j

where \vec{r}_{ij} is the normalised vector from \bar{x}_i to \bar{x}_j, and \vec{N}_i is the unit normal to element i at point \bar{x}_i (and vice versa for switching i and j).
```

RENDERING

- The resolution of the linear system gives the radiosities $B_i(\lambda)$ or B_i^R, B_i^G, B_i^B .
- Fix the view parameters
- Trace a ray from the viewpoint toward each pixel
- \bullet P: intersection point on a patch i.
- B_P is calculated by linear interpolation.

$$L_{pixel} = \frac{B_P}{\pi}$$

SOLVING THE SYSTEM

• Linear system :

$$KB = E, K_{ij} = -\rho_i F_{ij}, F_{ii} = 1 - \rho_i F_{ij}$$

• Solved by iterative methods:

$$B^{(k+1)} = f(B^{(k)}, B^{(k-1)}, ..., B^{(0)})$$

- $B^{(0)}$: initial guess
- Residual: r = E KB
- If $r^{(k)} \approx 0$ then $B^{(k)}$ is a good estimate.

Jacobi Method

$$\sum_{j} K_{ij} B_j = E_i \Rightarrow K_{ii} B_i = E_i - \sum_{j \neq i} K_{ij} B_j$$

 $\bullet \parallel r \parallel \leq \epsilon$

while not convergence do

for all i do
$$B_i^{(k+1)} = (E_i/K_{ii}) - \sum_{j \neq i} K_{ij} (B_j^{(k)}/K_{ii})$$

Gauss Siedel method Method

• Iteration

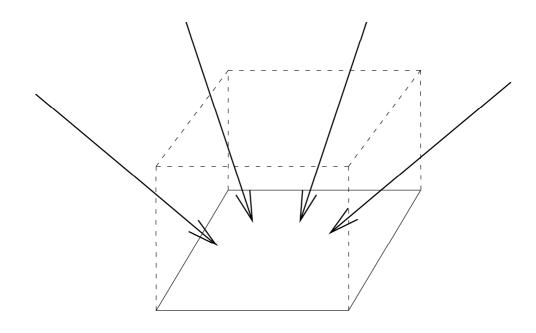
$$B_i^{(k+1)} = \frac{E_i}{K_{ii}} - \sum_{j=1}^{i-1} K_{ij} \frac{B_j^{(k+1)}}{K_{ii}} - \sum_{j=I+1}^n K_{ij} \frac{B_j^{(k)}}{K_{ii}}$$

$$\bullet \parallel B^{k+1} - B^k \parallel_{\infty} \le \epsilon$$

for all
$$i$$
 do
$$B_i = E_i;$$
while not convergence do
for all i do
$$B_i = (E_i/K_{ii}) + \sum_{j=1, j \neq i}^n B_j K_{ij}/K_{ii};$$

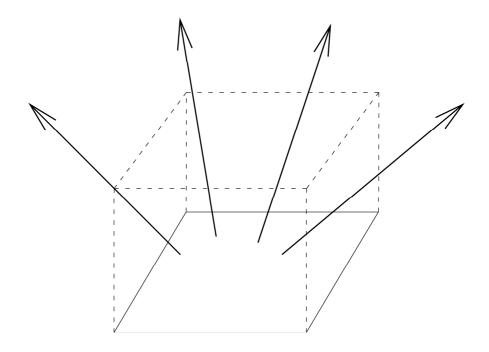
SOLVING THE SYSTEM

Complete solution



- Calculate the matrix system: memory and time complexities of $O(n^2)$.
- Resolution with Gauss-Siedel method $(K_{ij} = -\rho_i F_{ij}, K_{ii} = 1).$
- At each step, illumination from all other patches is gathered into a single receiving patch.
- The very large memory required for the storage of these form-factors limits the radiosity algorithm practically, except for hierarchical radiosity.

SOLVING THE SYSTEM:Progressive solution



- At each step, the illumination due to a single patch is distributed to all other patches within the scene. $B_{i \leftarrow j} = \rho_j B_i F_{ji} \frac{A_i}{A_j}$
- In the first steps, the light source patches are chosen to shoot their energy.
- The subsequent steps will select secondary sources, starting with those surfaces that receive the most light directly from the light sources, and so on.
- Each step increases the accuracy of the result that can be displayed.

SOLVING THE SYSTEM : Progressive solution

- The convergence criterion is met if $\|\Delta B \cdot A\|_{\infty}$ is below a certain threshold.
- This threshold could be a certain percentage of the sum of the total fluxes of the light sources.
- $\Delta B \cdot A$ is the vector of unshot fluxes.

```
for all i do

for each wavelength do

\Delta B_i = E_i;

while not convergence do

j = \text{patch-of-max-delta-flux}();

for all i do

for each wavelength do

\Delta Rad = \rho_i \Delta B_j F_{ji} \frac{A_j}{A_i};

\Delta B_i = \Delta B_i + \Delta Rad;

B_i = B_i + \Delta Rad;

\Delta B_j = 0;
```

Pseudo code for shooting

Progressive solution: Convergence and Ambient term

- After convergence, some residual fluxes remain unshot.
- Approximate them by an ambiant term.
- Ambiant term : $B_{ambiant} = R \Sigma_{j=1}^{N} \Delta B_{j} F_{*j}$
- Where F_{*j} represent the contribution of patch i to the others, and R characterises the mutiple interreflections.
- Expressions :

$$F_{*j} = \frac{A_j}{\sum_{k=1}^{N} A_k}$$

$$R = 1 + \rho_{ave} + \rho_{ave}^2 + \rho_{ave}^3 + \dots = \frac{1}{1 - \rho_{ave}}$$

 \bullet ρ_{ave} is the average reflectivity of the scene objects :

$$\rho_{moyen} = \frac{\sum_{k=1}^{N} \rho_k A_k}{\sum_{k=1}^{N} A_k}$$

• Updated radiosity: $B_i = B_i + \rho_i B_{ambiant}$.

RADIOSITY: Texture Mapping

- Distribution of the reflectivities for each wavelength over a textured patch.
- Calculate the average reflectivity ρ_{ave} for each textured patch, and compute the radiosity solution.
- Obtention of radiosities B_{ave} .
- Take into account the texture values only at the rendering step.
- The radiosity for a pixel is then:

$$B_{pixel} = B_{ave} \times \frac{\rho_{pixel}}{\rho_{ave}}$$

• ρ_{pixel} is the reflectivity of the scene point seen by the observer through the pixel.

HIERARCHICAL RADIOSITY

- Method to mesh the surfaces so that the radiosity be constant over each nesh.
- Avoids unuseful finer meshing
- Reduce the number of form factor calculations
- Subdivide the surfaces adaptively according to a criterion.
- A surface : a hierarchy of surface elements.
- Leaf = element, Node = group of elements.
- Interaction between two nodes A and B of different levels.
- \bullet Interaction if A and B exchange constant energy.
- One form factor for each interaction.
- Advantage: compute form factors not for each pair of leaf nodes but for each interaction.
- Reduction of computation and memory storage.
- Link(A, B): when the two nodes interact.

HIERARCHICAL RADIOSITY: Data Structures

```
struct Quadnode {
     float B_q[]; /* gathering radiosity at sample \lambda's */
     float B_s[]; /* shooting radiosity at sample \lambda's*/
     float E[\ ]; /* self emittance at sample \lambda's*/
     float area;
     float \rho[] /* reflectivity at sample \lambda's;
     struct Quadnode** children; /* pointer to list of
                      four children*/
     struct Linknode* L; /* first gathering link of node */
}
struct Linknode {
     Quadnode* q; /* gathering node */
     Quadnode* p; /* shooting node */
     float F_{qp}; /* form factor from q to p */
     struct Linknode* next; /* next gathering link of node q
*/
};
```

HIERARCHICAL RADIOSITY: Hierarchy and Interactions

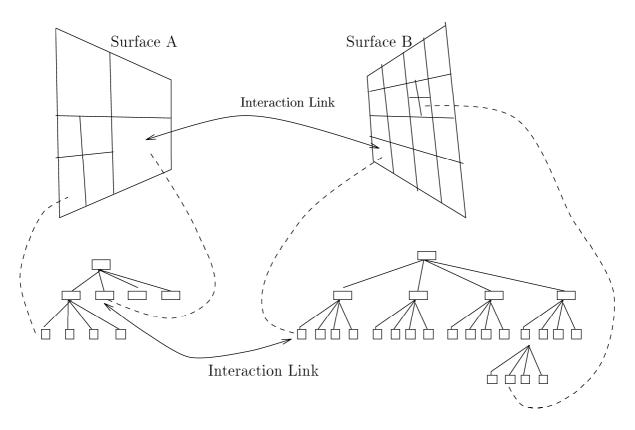


Figure 2: Hierarchy and interactions

HIERARCHICAL RADIOSITY: Refinement

- Subidivide a surface with respect to the others.
- A surface may be finely subdivided with respect to a surface A, and coarsely with respect to another surface B.

```
Refine(Quadnode *p, Quadnode *q, float F_{\epsilon})

{
    Quadnode which, r;
    if (oracle(p, q, F_{\epsilon})) Link(p, q)
    else
        which = Subdiv(p, q);
    if (which == q)
        for (each node r of q) Refine(p, r, F_{\epsilon})
    else if (which == p)
        for (each node r of p)Refine(r, q, F_{\epsilon})
    else
        Link(p,q);
}
```

Refine pseudo code.

Solving the Hierarchical System

- Gather: gathers energy over each link at each receiving node. Jacobi iteration.
- **PushPull** pushes the gathered radiosity down to the children of each receiving node, and pulls the results back up the quadtrees by area averaging (decomposition and reconstruction).
- Convergence: maximum change below a threshold.

Solving the Hierarchical System: Gathering

```
GatherRad(Quadnode *p) {
    Quadnode *q;
    Link *L;

p \to B_g = 0;
    for(each gathering link L of p)
        /* gather energy across link */
        p \to B_g += p \to \rho * L \to F_{pq} * L \to q \to B_s;
    for(each child node r of p)
        GatherRad(r);
}
```

Solving the Hierarchical System: PushPull

```
PushPullRad(Quadnode *p, float B_{down})
{
	float B_{up}, B_{tmp};
	if(p \to children == \text{NULL}) /* p is a leaf */
	B_{up} = p \to E + p \to B_g + B_{down};
	else
	B_{up} = 0;
	for (each child node r or p)
	B_{tmp} = \text{PushPullRad}(r, p \to B_g + B_{down})
	B_{up} += B_{tmp} * \frac{r \to area}{p \to area};
	p \to B_s = B_{up};
	\text{return}(B_{up});
}
```

The Oracle

- Takes the decision to link or not two nodes.
- A link is built if F_{pq} is small enough to consider the energy contribution of p to q as small, which amounts to say that the radiosity of q due to p can be considered constant over q (**Oracle1**).
- F_{pq} is estimated as:

$$F_{pq} \approx \frac{\cos \theta}{\pi} \Omega_q,$$

where Ω_q is the solid angle whose appex is the center of p and subtented by a disk surrounding q.

A better Oracle

- Oracle1 uses a geometric subdivision criterion.
- This may results in a large number of fine elements.
- More subtle to use a criterion based on the amount of energy transferred between two nodes.
- If $F_{pq} \cdot B_q \cdot A_q \leq BF_{\epsilon}$ then a link is established.
- Since the radiosities are not known a priori, the refinement algorithm proceeds adaptively by using another oracle **Oracle2**.

```
float \mathbf{Oracle2}(\mathbf{Linknode} *L, \text{ float } BF_{\epsilon}) {
    Quadnode *p = L \rightarrow p; /* shooter */
    Quadnode *q = L \rightarrow q; /* receiver */
    if (p \rightarrow area < A_{\epsilon} \text{ and } q \rightarrow area < A_{\epsilon})
        return(FALSE);
    if (p \rightarrow B_s == 0.0)
        return(FALSE);
    if ((p \rightarrow B_s * p \rightarrow Area * L \rightarrow F_{pq}) < BF_{\epsilon})
        return(FALSE);
    else
        return(TRUE);
}
```

The Hierarchical Radiosity Algorithm

- Refine uses Oracle1() and establishes links at the highest levels unless the shooting surface is a light source.
- Most of these links are built even though the shooting radiosities of most the surfaces are zero.
- These links will be refined in the second pass through **RefineLink** which use **Oracle2**

The Algorithm

```
HierachicalRad(float BF_{\epsilon})
\left\{ \right.
    Quadnode *p, *q;
    Link *L;
    int Done = FALSE;
    for (all surfaces p) p \to B_s = p \to E;
    for (each pair of surfaces p, q)
        Refine(p, q, BF_{\epsilon});
    while (not Done){
        Done = TRUE;
        SolveSystem();
        for (all links L)
             /* RefineLink returns FALSE if any
               subdivision occurs */
             if (RefineLink(L, BF_{\epsilon}) == FALSE)
                 Done = FALSE;
    }
```

The Hierarchical Radiosity Algorithm: RefineLink

```
int \mathbf{RefineLink}(\mathrm{Linknode} *L, \mathrm{float}\ BF_{\epsilon}) {
    int no\_subdivision = \mathrm{TRUE};
    Quadnode *p = L \to p; /* shooter */
    Quadnode *q = L \to q; /* receiver */

    if (\mathbf{Oracle2}(L, BF_{\epsilon}))
    no\_subdivision = \mathrm{FALSE};
    which = \mathbf{Subdiv}(p, q);
    \mathbf{DeleteLink}(L);
    if (which == q)
    for (each child node r of q) \mathbf{Link}(p, r);
    else
    for (each child node r of p) \mathbf{Link}(r, q);
    return (no\_subdivision);
}
```

GENERAL GLOBAL SOLUTION

Several approaches

• One-pass:

- Illumination computations performed independently of the view point.
- Fast rendering for different viewpoints.
- Place a small relecting surface at the observer, or a non reflecting surface covering the whole virtual screen.
- In this last case, additionnal computations are limited to the L_{ij} between the screen surface and the visible surfaces of the screen.
- View independence.
- Drawbacks: large memory to store data, aliasing defects due to sharp variations of specular reflections and transmissions.
- IMMEL et al. (Siggraph, 1986), MUDUR et al. (Visualization and Computer Animation, 1990), LE SAECK et al. (Eurographics Workshop, 1990), SHAO et al. (Siggraph, 1988).

• Two-pass:

- Global iffuse and global specular components from reflection and transmission are calculated separately.
- Specular component is evaluated once the global one has been computed.
- Global diffuse component: solution of a system of linear equations (similar to Radiosity)
- Extended form-factors.
- Global specular component: ray tracing, distributed ray tracing, Monte-Carlo.
- Global diffuse component is view independent.
- For different viewpoints, only the specular component has to be evaluated.
- SILLION et al. (Siggraph, 1989, 1991), WAL-LACE et al. (Siggraph, 1987, 1989), HECBERT (Siggraph, 1990)

• Multi-pass:

- Several passes.
- Caustic effects.
- Monte-Carlo, path tracing, ray tracing, progressive radiosity
- SHIRLEY (Graphics Interface, 1990), CHEN et al.(Siggraph, 1991)

Visual perception

- The human eye converts luminance into a visual sensation, called *brightness*.
- The range of visible luminance is 10^{-6} to 5.10^4 cd/m².
- The visual sensation is related to the luminance, but is not linear, and depends on the ambiant level of illumination.
- Approximation: the law of sensitivity is logarithmic (Weber' law).
- Why the sensitivity varies? Because:
 - the size of the iris varies with luminance,
 - then the sensitivity of the retina is modified.

Visual perception

- Take into account the sensitivity of the eye: contexts of a real scene as seen by the eye, and display device.
- Difficulty: find a function (Tumblin and Rushmeier) relating the luminance of the real world to the values to be displayed on a monitor.
- These values depend on: characteristics of the display device, and ilumination of the room containing this device.
- Overcome this difficulty: scale the image so that it fits in the color range of the monitor, and correct the non-linearities of the monitor.

Color clipping and scaling

- Transformation of spectral radiances into a color space (XYZ, then RGB).
- Range of RGB values very important (highlights due to specular reflections).
- Daylight is also source of important radiance variations.
- Sometimes, negative values which cannot be displayed.
- Scale and clip these components to obtain images with a maximum dynamic and a minimum loss.

Color scaling

- Technique used by lighting engineers.
- Consider a scene lit by artificial light sources.
- Known data: sum the luminous fluxes emitted by these sources.
- Which gives Φ_{total} .
- Assume that the total emitted flux reaches only the floor of the scene.
- Assume also that the floor is diffuse and has an average reflectivity ρ_{ave} .
- Then, the approximated average radiance of the floor is simply:

$$L_{ave} = \frac{\rho_{ave} \times \Phi_{total}}{\pi floor_area}.$$

• Assume that the maximum radiance L^{max} to be displayed is fixed to twice L_{ave} , then the range for RGB components is

$$[0, 2L_R^{max}] \times [0, 2L_G^{max}] \times [0, 2L_B^{max}].$$

- Assume that the spectral distribution $L^{max}(\lambda)$ associated with L^{max} is equal to $C\Phi(\lambda)$ where C is a constant and $\Phi(\lambda) = 1 \,\forall \lambda$.
- Then:

$$\begin{split} L_X^{max} &= K \int C \, \bar{x}(\lambda) \, d\lambda, \, L_Y^{max} = K \int C \, \bar{y}(\lambda) \, d\lambda, \\ L_Z^{max} &= K \int C \, \bar{z}(\lambda) \, d\lambda. \end{split}$$

• This yields approximatively

$$L_X^{max} = L_Y^{max} = L_Z^{max} = KC$$

- .
- These XYZ components are transformed to RGB components to give $L_R^{max}, L_G^{max}, L_B^{max}$.
- Let MAXDISPLAY be equal to twice the maximum value of these three components.

• The final RGB components of the calculated radiances are computed by:

```
for each pixel
{
    if at lest one component is negative or bigger than
MAXDISPLAY
    then clip it
/* scale */
    for each component C of the pixel
        Cdisplay = (C/MAXDISPLAY)*255
}
```

Color clipping

- not possible to display all the calculated colors on a monitor, some of them beeing out of the gamut of the monitor, and others exceeding its range.
- Different methods of clipping (HALL89).
- Scale and clip the entire image until there are no luminances too high for display.
- Or, maintain the chromaticity and scale the luminance of the offending color.
- Or, maintain the dominant hue and luminance and desaturates the color.
- Or, clamp any color component exceeding 1 to 1.
- No method gives the best results in any case.
- Thus, set the negative value to 0, and values bigger than MAXDISPLAY to MAXDISPLAY.
- But choose an appropriate MAXDISPLAY.

Gamma correction

- The luminance L (Y component) produced by the phosphors of a monitor is not proportional to the input signals R, G, B.
- Non linear response: $L = kI^{\gamma}$.
- Where γ is a parameter depending on the monitor, and is about 2.3 for the three RGB channels of typical rasters.
- Assume that this law is valid for each RGB component.
- Then replace, for example, the calculated component R by $R^{1/\gamma}$.

Conditions for display

- Make sure that the room containing the display device is very dark, in order to avoid reflections on the screen.
- Then calibrate the monitor.
- To do this, display a totally black picture and set the brightness control so that you are just under the perception level.
- This setup must not be modified until the viewing conditions are changed.